MTH 437/537

HW #6

The general Robertson-Walker (isotropic, homogeneous) cosmological model has line element

$$ds^{2} = -dt^{2} + a(t)^{2} \left( \frac{dr^{2}}{1 - kr^{2}} + r^{2} d\theta^{2} + r^{2} \sin^{2} \theta d\phi^{2} \right)$$

with k = -1, 0, 1 depending on whether the spatial cross-sections are hyperbolic, flat, or spherical, respectively. The only nonzero components of the Einstein tensor are

$$G^{t}_{t} = -3 \frac{\dot{a}^{2} + k}{a^{2}}$$
;  $G^{r}_{r} = G^{\theta}_{\theta} = G^{\phi}_{\phi} = -\frac{2a\ddot{a} + \dot{a}^{2} + k}{a^{2}}$ 

(You do not need to verify this.)

## 1. CURVATURE SCALARS

The Einstein tensor is related to the Ricci tensor by

$$G^i{}_j = R^i{}_j - \frac{1}{2}\,\delta^i{}_j\,R$$

(a) Find an expression for the "Einstein scalar"  $G = G^{i}{}_{i}$  in terms of the Ricci scalar  $R = R^{i}{}_{i}$ .

(b) Determine the Ricci scalar for the Robertson-Walker model above.

(c) (**Optional!**) Derive at least one component of the Einstein tensor above for the case k = 0.