## Cosmology Problem Set #1

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The following exercises are due Friday, February 2.

- 1. An observer on earth watches a spaceship traveling with a speed of  $\beta = 0.9$ . He uses a telescope to watch a clock ticking on board the ship. How fast does he *see* the clock tick relative to the proper time. (Hint: Make allowances for the travel time of the light signal.)
- 2. Here is an exercise in using tensor notation. The scalar product between the four-vectors x and y in special relativity is

$$x \cdot y = x^{\mu} \eta_{\mu\nu} y^{\nu}$$

Lorentz transformations, which we write

$$x^{\prime \mu} = [\boldsymbol{L}]^{\mu}_{\ \nu} x^{\nu},$$

leave this product invariant, and do not change  $\eta$ . Use this to prove that

$$\eta_{\gamma\delta} = [\boldsymbol{L}]^{\alpha}_{\ \gamma}\eta_{\alpha\beta}[\boldsymbol{L}]^{\beta}_{\ \delta}$$

We sometimes use this as a definition of the Lorentz transformation.

- 3. Exercise (2.3) in the text. I realize that a brief solution is given in the back of the book. I would like you to write it out carefully so that you understand it.
- 4. Here is another way to derive the formula for the Doppler shift. It make use of the fact that the  $k^{\alpha} = (\omega/c, \mathbf{k})$  is a four-vector. Apply the Lorentz transformation to this quantity to show that the frequency observed in the laboratory for a source moving in the  $\hat{\mathbf{x}}$  direction with velocity  $v = \beta c$ for arbitrary  $\mathbf{k}$  is given by the formula

$$\frac{\omega}{\omega_0} = \frac{1}{\gamma(1 - \beta\cos\theta)}$$

where  $\omega_0$  is the frequency in the rest frame of the source, and  $\theta$  is the angle of  $\mathbf{k}$  relative to  $\hat{\mathbf{x}}$  in the laboratory frame of reference.

Cosmology Sit # 1



To make it easy assume the first tick of the  
clock occurs at 
$$t=t'=0$$
 when the arigins. We  
can the identity 3 point events:  
 $\pm 1$  the clock ticks  $\chi_1 = \chi_1' = t_1 = t_1' = 0$   
 $\pm 2$  the clock tecks a second time  
 $t_2' = \Delta t$   $\chi_2' = 0$   $\chi_2 = 2 t_2$   
 $\pm 3$  the deght from  $\pm 2$  reaches the observer  
 $t_3 = 7$   $\chi_3 = 0$ 

Because of the travel time of light 
$$t_3 = t_2 + x_2/c$$
  
 $t_3 = t_2 (1+b/c)$ 

The times are related by 
$$t = t (t' + \beta x/c)$$
  
In this case

$$t_2 = f(t_2' + px_2'/c) = f \Delta t$$

$$\frac{t_3}{At} = \frac{(1+B)}{\sqrt{1-B^2}} = 4.36$$

So if the clock ticks once a second, the observa will see it clear every 4-36 sec.

2. x'. q'= x' " huv q' " = L " x x y w L p y B = x y = x Y dp y B This must be true for all ocd and yB, so La Vue Lop = Uxp 3. (a) chain rule:  $\frac{\partial}{\partial x'^{\mu}} = \frac{\partial x'^{\nu}}{\partial x'^{\mu}} \frac{\partial}{\partial x^{\nu}}$ but  $\mathcal{D}C' = \left[ L^{-1} \right]^{\mathcal{D}} \mathcal{D}C''$  $\sigma_{\perp} = \begin{bmatrix} \Delta \chi^{\nu} \\ \Theta \chi^{\prime} M \end{bmatrix} = \begin{bmatrix} L^{-1} \end{bmatrix}^{\nu} M$ and dra = [L-1]~ u dr  $\alpha \downarrow \Theta_{\mathcal{M}} \equiv \underbrace{\Theta}_{\mathcal{M}} \qquad \Theta_{\mathcal{M}} = \begin{bmatrix} L^{-1} \end{bmatrix}^{\mathcal{D}}_{\mathcal{M}} \qquad \Theta_{\mathcal{D}}$ (b) assume  $\Theta_{xe}' = [I]_{u} \vee \partial_{v}$  where I is some unracion matrix. Then  $\frac{\partial x}{\partial r''} = \partial_{\mu} x'' = [I]_{\mu} \partial_{\alpha} [L]_{\beta} x^{\beta}$ = [I] d [L] p 8d

 $= [I]_{u} \land [L] \land = S_{u}$ So I = 1-1 as promised. Hd= (W/c, F) 4.  $K' = [L] B K^{B}$  $\begin{array}{c|c} \omega' c & \lambda & -\beta \lambda & 0 & 0 & | \omega | c \\ H_{\chi} & -\beta \lambda & + & 0 & 0 & | H_{\chi} \\ H_{\chi} & 0 & 0 & 1 & 0 & | H_{\chi} \\ H_{\chi} & 0 & 0 & 0 & 1 & | H_{\chi} \\ \end{array}$  $\frac{\omega}{c} = t \left( \frac{\omega}{c} - \beta K_{z} \right)$ but Hoc = K COSG = W COSG  $90 W' = + W (1 - \beta \cos \alpha)$