

Rotations in the state space

$$|\psi\rangle \Rightarrow |\psi'\rangle = \mathcal{D}(R) |\psi\rangle$$

↑

Drehung ← rotation (German)

Need to preserve $\langle \psi | \psi \rangle = \langle \psi' | \psi' \rangle$

⇓
 $\mathcal{D}(R)$ must be unitary!

$$\mathcal{D}(R) \mathcal{D}^\dagger(R) = I$$

⇓
 $\mathcal{D}^{-1}(R)$

⇓ Recall Lectures #10-11
of Phys 651

From Phys 651 ⇒ unitary transformations

$$U_\epsilon = 1 - i G \epsilon$$

angular momentum $G = \frac{J_z}{\hbar}$,

$$\epsilon = d\psi$$

↓
rotation
around z-axis
by $d\psi$

← generator

$$G = \frac{H}{\hbar}, \epsilon = dt$$

↓
time evolution
by dt

$$G = \frac{p_x}{\hbar}, \epsilon = dx$$

↓
translation
by dx

⇒

$$\mathcal{D}(\vec{n}, d\psi) = 1 - i \left(\frac{\vec{J} \cdot \vec{n}}{\hbar} \right) d\psi$$

rotation around \vec{n} by $d\psi$

or, in the case of finite rotations \Rightarrow

$$\mathcal{D}(\vec{n}, \psi) = \left(1 - i \frac{\vec{J} \cdot \vec{n}}{\hbar} d\psi \right) \left(1 - i \frac{\vec{J} \cdot \vec{n}}{\hbar} d\psi \right) \dots$$

$$= \lim_{N \rightarrow \infty} \left[1 - i \frac{\vec{J} \cdot \vec{n}}{\hbar} \frac{\psi}{N} \right]^N = \exp\left(-\frac{i}{\hbar} (\vec{J} \cdot \vec{n}) \psi\right)$$

N times

What happens to observables upon rotation? \Rightarrow
Lecture #10 of Phys 657 \Rightarrow

$$A' = \mathcal{D}(R) A \mathcal{D}^\dagger(R), \quad A = \mathcal{D}^\dagger(R) A' \mathcal{D}(R)$$

Let's apply infinitesimal rotation and see how A changes:

$$A' = \left(1 - i \frac{\vec{J} \cdot \vec{n}}{\hbar} d\psi \right) A \left(1 + i \frac{\vec{J} \cdot \vec{n}}{\hbar} d\psi \right) =$$

$$= A - \frac{i}{\hbar} d\psi [\vec{J} \cdot \vec{n}, A] = A'$$

neglect $O(d\psi^2)$

Definition; an observable is scalar if $A=A'$ ⁽³⁾
i.e. $[\vec{J} \cdot \vec{n}, A] = 0$

If an observable commutes with $\vec{J} \cdot \vec{n}$
(i.e. generator of rotation about \vec{n} -axis)
it does not change upon rotation about \vec{n} .

What if $A=H$ \nwarrow Hamiltonian of an isolated system

Consider an isolated system in a state $|\Psi(t_0)\rangle$. Under rotation, the state transforms into $|\Psi'(t_0)\rangle = \mathcal{D}(R) |\Psi(t_0)\rangle$

At a time t_0+dt , the state is

$$|\Psi'(t_0+dt)\rangle = |\Psi'(t_0)\rangle + dt \cdot \left. \frac{d|\Psi'\rangle}{dt} \right|_{t_0}$$
$$= |\Psi'(t_0)\rangle + \frac{dt}{i\hbar} H |\Psi'(t_0)\rangle = \mathcal{D}(R) |\Psi(t_0+dt)\rangle$$

\uparrow
 $i\hbar \frac{\partial \Psi'}{\partial t} = H \Psi'$ (9.1)

If we had not performed rotation \Rightarrow the state at t_0+dt would be:

$$|\Psi(t_0+dt)\rangle = |\Psi(t_0)\rangle + \frac{dt}{i\hbar} H |\Psi(t_0)\rangle$$

Now rotate \Rightarrow apply $\mathcal{D}(R)$ to $\Psi \Rightarrow$

$$\underbrace{D(R) |\Psi(t_0 + dt)\rangle}_{\parallel} = \underbrace{D(R) |\Psi(t_0)\rangle}_{\parallel} + \quad (4)$$

$$+ \frac{dt}{i\hbar} D(R) H |\Psi(t_0)\rangle \quad (9.2)$$

Compare (9.1) and (9.2) \Rightarrow

$$H \underbrace{|\Psi'(t_0)\rangle}_{D(R) |\Psi(t_0)\rangle} = D(R) H |\Psi(t_0)\rangle \Rightarrow \underbrace{[H, D(R)] = 0}$$

$$\underbrace{H \text{ is a scalar observable}}_{\uparrow} \Leftrightarrow \underbrace{[H, \vec{J} \cdot \vec{n}] = 0}$$

↑
does not change upon rotation

From Phys 651 \Rightarrow time evolution of the mean value of an observable $A \Rightarrow$

$$\frac{d\langle A \rangle}{dt} = \left\langle \frac{\partial A}{\partial t} \right\rangle + \frac{1}{i\hbar} \langle [A, H] \rangle$$

$$\text{If } A \equiv \vec{J} \Rightarrow \frac{d\langle \vec{J} \rangle}{dt} = \frac{1}{i\hbar} \langle [\vec{J}, H] \rangle = 0$$

total angular momentum of an isolated system is a constant of motion: conservation of angular momentum is a consequence of rotational invariance. (5)

Recall Lecture # ~~8~~ (on geometric rotations)

$$[R_x(\epsilon), R_y(\epsilon)] = R_z(\epsilon^2) - 1$$

$$\Downarrow \text{QM} \Leftrightarrow \mathcal{D}(\vec{n}, d\psi)$$

$$\left(1 - \frac{i}{\hbar} J_x \epsilon - \frac{J_x^2 \epsilon^2}{2\hbar^2}\right) \left(1 - i \frac{J_y}{\hbar} \epsilon - \frac{J_y^2 \epsilon^2}{2\hbar^2}\right) -$$

↑
keep $O(\epsilon^2)$

$$- (\text{same, } y \leftrightarrow x) = \cancel{1} - \frac{i}{\hbar} J_z \epsilon^2 - \cancel{1} \xRightarrow{\text{WS!}} \text{HW!} \leftarrow \text{details}$$

\Downarrow

generalizing to other axes

$$\Leftrightarrow [J_x, J_y] = i\hbar J_z$$

$$[J_i, J_j] = i\hbar \epsilon_{ijk} J_k$$

↑ Levi-Civita tensor

$$\epsilon_{ijk} = \begin{cases} +1 & \leftarrow \text{even permutation} \\ -1 & \leftarrow \text{odd permutation} \\ 0 & \leftarrow \text{otherwise} \end{cases}$$

So, how would group theory help? (6)

We could have avoided all derivations ~~above!~~

Noether's theorem:

For every continuous symmetry there exists a corresponding conservation law, i.e. there exists a conserved observable



Group \Rightarrow Unitary representation $U = e^{i\epsilon G}$

Wigner theorem:

for every symmetry group there is a unitary representation

$$U = e^{i\epsilon G} \quad \leftarrow \text{generator}$$

$$|\psi'\rangle = U|\psi\rangle$$



$$H' = U(\epsilon) H U^{-1}(\epsilon)$$

$$= H$$

↑ Hamiltonian is invariant under this unitary transformation



$$[H, G] = 0 \Rightarrow$$

G is a constant of the motion

$$[G_i, G_k] = C_{ikl} G_l$$

generators of a symm. group are conserved observables whose commutation relations are uniquely determined by group structure. \uparrow depends on G \Rightarrow (if $G \equiv J \Rightarrow C_{ikl} \equiv \epsilon_{ikl}$
 $G \equiv P \Rightarrow C_{ikl} = 0$)