

so that:

$$\int_{-\infty}^{+\infty} |\Psi(x)|^2 dx = 1$$

We say that the $\Psi(x)$ function must be normalized - and the above equation expresses this "normalization requirement" mathematically.

APPLICATIONS OF TISE:

We will now discuss the application of Schrödinger Equation to several simple quantum-mechanical situations (but "simple" does not mean that they are not important!!!).

FREE PARTICLE:

The simplest of all conceivable quantum mechanical systems is certainly a free particle moving in the x direction, with no forces whatsoever acting on it. $F=0$ means that the potential energy is constant: $U(x) = \text{const.}$

As you remember, in the case of potential energy we can arbitrarily choose where $U = 0$.

Question: when we talk about the potential energy due to Earth gravity, we often choose one special point to be that with $U = 0$. Where is this point located?

So, for the "free-particle" we can set $U = 0$ and not worry any more about it!

Then, the TISE becomes:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} = E \psi$$

or $\frac{d^2 \psi}{dx^2} = -k^2 \psi$ with $k^2 = \frac{2mE}{\hbar^2}$

This is a well-known equation, with a general solution:

$$\psi = A \sin(kx) + B \cos(kx)$$

where A and B are arbitrary constants.

This is an equation of a wave with $\lambda = \frac{k}{2\pi}$

The equation is satisfied for any value of k meaning that the energy is:

$$k^2 = \frac{2mE}{\hbar^2} \Rightarrow E = \frac{\hbar^2 k^2}{2m}$$

So, the free-particle equation describes essentially the motion of ~~any~~ a particle with any energy E .

Now, note that for non-relativistic particles:

$$\left. \begin{aligned} E = K &= \frac{mV^2}{2} = \frac{m^2 V^2}{2m} = \frac{p^2}{2m} \\ \text{and} \quad E &= \frac{\hbar^2 k^2}{2m} \end{aligned} \right\} \Rightarrow p = \hbar k$$

$$\text{But } \hbar = \frac{h}{2\pi} \text{ and } k = \frac{2\pi}{\lambda}$$

$$\text{So: } p = \hbar k = \frac{h}{2\pi} \cdot \frac{2\pi}{\lambda} = \frac{h}{\lambda}, \text{ or } \lambda = \frac{h}{p}$$

The latter is simply the de Broglie wavelength.

So, the solution produced simply a de Broglie wave, which is not surprising!

But there is a small problem---

(not mentioned in the book!)

We said that $\psi(x)$ must be normalized!

And

$$\int_{-\infty}^{+\infty} |A \sin(kx) + B \cos(kx)|^2 dx$$

$$= \int_{-\infty}^{+\infty} [A^2 \sin^2(kx) + B \cos^2(kx) + AB \sin(2kx)] dx = \infty !!!$$

So, did our very first encounter with the TISE produce an illegal solution?

Fortunately, ψ can be rescued!

Since $\psi(x) = A \sin(kx) + B \cos(kx)$ for any k is a "good" solution, any sum of those solutions ~~is also~~ with different k values is also a good solution. And by adding many such solutions, we can obtain a wave packet!

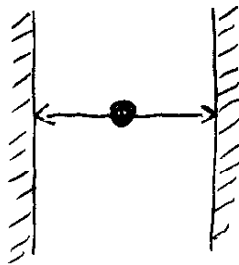
ψ_1  $\int_{-\infty}^{+\infty} |\psi|^2 dx = \infty$

infinite wave

ψ_2  finite wave-packet
(Gaussian) $\int_{-\infty}^{+\infty} |\psi|^2 dx = \text{finite value.}$

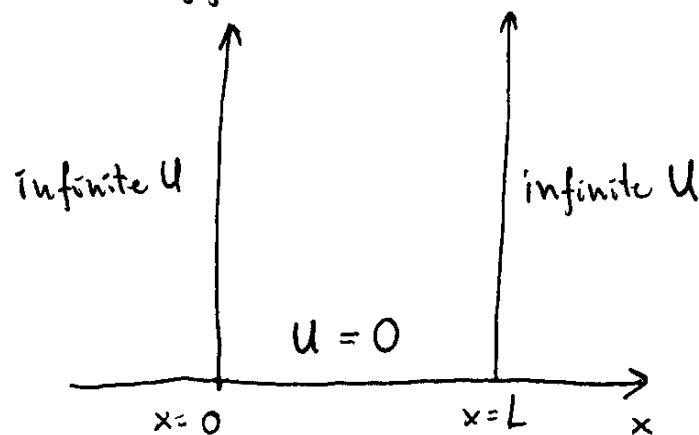
2. PARTICLE IN A BOX

Meaning: particle between two walls. The collisions with each wall are elastic, i.e., the particle is bounced back with no energy loss.



This is again a simple problem - but it has many important implications, and is used in many "serious" physical models.

We describe this situation in terms of potential energy:



The particle cannot be in the regions $x < 0$ and $x > L$, so $\psi(x < 0) = 0$ and $\psi(x > L) = 0$

Only for $0 \leq x \leq L$ the wavefunction can be $\psi(x) \neq 0$.

The Schrödinger Eq. for the region between the walls is:

$$-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} = E\psi$$

And, as was said before, the solution for ψ is:

$$\psi = A \sin(kx) + B \cos(kx) \quad \text{with } k^2 = \frac{2mE}{\hbar^2}$$

However, now there are some restrictions imposed on ψ . First, $\psi(x)$ must be a continuous function. For $x < 0$, as we said, $\psi = 0$.

So, the ~~solution~~ "inside" solution also must be zero at $x=0$. And $\cos(kx)$ is one for $x=0$.

So, we conclude that B must be zero.

This leaves us with:

$$\psi = A \sin(kx)$$

But, again, $\psi = 0$ for all $x > L$, so to assure the continuity of ψ , it must be:

$$\psi(L) = A \sin(kL) = 0$$

Meaning that:

$$kL = 0, \pi, 2\pi, 3\pi, \dots, n\pi, \dots$$

$kL = 0$ would mean $\psi = 0$ everywhere, or no particle. So, it has to be rejected.

Other kL values are O.K. So we get:

$$k = \frac{n\pi}{L} \quad \text{with } n = 1, 2, 3, \dots$$

and:

$$\psi = A \sin\left(\frac{n\pi}{L}x\right)$$

Still, we have to determine the value of A .

Here, the normalization condition becomes helpful:

$$\int_{-\infty}^{+\infty} |\psi(x)| dx = 1 \quad \xrightarrow{\text{in the present case}} \quad \int_0^L A^2 \sin^2\left(\frac{n\pi}{L}x\right) dx = 1$$

Use a dummy variable $z = \frac{n\pi}{L}x$: $dx = \frac{L}{n\pi} dz$

$$\int_0^L A^2 \sin^2\left(\frac{n\pi}{L}x\right) dx = A^2 \frac{L}{n\pi} \int_0^{n\pi} \sin^2(z) dz = A^2 \frac{L}{n\pi} \cdot \frac{n\pi}{2} = \frac{A^2 L}{2} = 1$$

so $A = \sqrt{\frac{2}{L}}$

Our wavefunction is then:

$$\psi = \begin{cases} \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L}x\right) & \text{for } 0 \leq x \leq L \\ 0 & \text{for } x < 0 \text{ and } x > L \end{cases}$$

But we still have to find the energy E .

This is easy:

$$k^2 = \frac{2mE}{\hbar^2} \quad \text{and} \quad k = \frac{n\pi}{L}$$

$$\text{So: } E_n = \frac{\hbar^2 \pi^2 n^2}{2mL^2} \quad \text{with } n = 1, 2, 3, \dots$$

The lowest allowed energy is for $n=1$:

$$E_1 = \frac{\hbar^2 \pi^2}{2mL^2}$$

The other possible energies can be simply written as:

$$E_n = n^2 E_1$$

$$\text{So } E_2 = 4E_1, \quad E_3 = 9E_1, \quad E_4 = 16E_1, \dots$$