Analysis of Methods for Dispersive Electromagnetics with Distributions of Parameters

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#### 2 Models and Methods

- Random Lorentz
- Polynomial Chaos
- FDTD

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# Stability

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# Background

Maxwell's Equations:

$$\frac{\partial \mathbf{D}}{\partial t} + \mathbf{J} = \nabla \times \mathbf{H}$$
$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$
$$\nabla \cdot \mathbf{D} = \rho$$
$$\nabla \cdot \mathbf{B} = 0$$

Constitutive Relations:

$$\mathbf{D} = \epsilon \mathbf{E} + \mathbf{P}$$
$$\mathbf{B} = \mu \mathbf{H} + \mathbf{M}$$
$$\mathbf{J} = \sigma \mathbf{E} + \mathbf{J}_s$$

Boundary Conditions:

$$\mathbf{E} \times \mathbf{n} = 0$$
, on  $(0, T) \times \partial D$ ,

Initial Conditions:

 $\mathbf{E}(0,\mathbf{x})=0,\quad \mathbf{H}(0,\mathbf{x})=0, \text{ in } \mathcal{D}.$ 

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We employ the physical assumption that electrons behave as damped harmonic oscillators,

$$m\ddot{x} + 2m\nu\dot{x} + m\omega_0^2 x = F_{driving}$$

The polarization is then defined as electron dipole moment density:

$$\ddot{P} + 2\nu \dot{P} + \omega_0^2 P = \epsilon_0 \omega_p^2 E$$

where  $\omega_0$  is the resonant frequency,  $\nu$  is a damping coefficient, and  $\omega_p$  is referred to as a plasma frequency defined by  $\omega_p^2 = (\epsilon_s - \epsilon_\infty)\omega_0^2$ .

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Taking a Fourier transform of  $D = \epsilon E + P$  and inserting the convolution form of the polarization model in for P, we get  $\hat{D}(\omega) = \epsilon_0 \epsilon(\omega) \hat{E}(\omega)$  where

$$\epsilon(\omega) = \epsilon_{\infty} + \frac{\omega_{\rho}^2}{\omega_0^2 - \omega^2 - i2\nu\omega}.$$

For multiple Lorentz poles, the complex permittivity includes a (weighted) sum of mechanisms:

$$\epsilon(\omega) = \epsilon_{\infty} + \sum_{i=1}^{N_p} rac{\omega_{p,i}^2}{\omega_{0,i}^2 - \omega^2 - i2
u_i\omega}.$$

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The multi-pole Lorentz model motivates a model with a continuum of Lorentz mechanisms, i.e., a distribution of dielectric parameters. We define a random polarization to be a function of a dielectric parameter treated as a random variable.

The random Lorentz model is

$$\ddot{\mathcal{P}} + 2\nu\dot{\mathcal{P}} + \omega_0^2\mathcal{P} = \epsilon_0\omega_\rho^2 E$$

with parameter  $\omega_0^2$  treated as a random variable with probability distribution F on the interval (a, b). The macroscopic polarization is taken to be the expected value of the random polarization,

$$P(t,z) = \int_a^b \mathcal{P}(t,z;\omega_0^2) \, dF(\omega_0^2).$$

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# **Random Polarization**



# Complex Permittivity with random $\omega_0^2$

Separate complex permittivity into real and imaginary parts ( $\epsilon = \epsilon_r + i\epsilon_i$ ):

$$\epsilon_r = \epsilon_{\infty} + \frac{\omega_p^2(\omega_0^2 - \omega^2)}{(\omega_0^2 - \omega^2)^2 + 4\nu^2\omega^2}$$
$$\epsilon_i = \frac{2\omega_p^2\nu\omega}{(\omega_0^2 - \omega^2)^2 + 4\nu^2\omega^2}.$$

Analytic integration is possible for uniform distribution:

$$\mathbb{E}[\epsilon_r] = \frac{1}{b-a} \int_a^b \epsilon_r d\omega_0^2 = \epsilon_\infty + \frac{\omega_\rho^2}{2(b-a)} \left( \ln(\omega_0^2)^2 - 2\omega_0^2 \omega^2 + \omega^4 + 4\nu^2 \omega^2 \right) \Big|_a^b$$

$$\mathbb{E}[\epsilon_i] = \frac{1}{b-a} \int_a^b \epsilon_i d\omega_0^2 = \frac{\omega_p^2}{(b-a)} \arctan\left(\frac{\omega^2 - \omega_0^2}{2\nu\omega}\right) \Big|_a^b$$

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# Saltwater Data



Figure 2: Fits for single-pole, saltwater data

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# Maxwell-Random Lorentz system

In a polydisperse Lorentz material, we have

$$\epsilon_0 \epsilon_\infty \frac{\partial \mathbf{E}}{\partial t} = \nabla \times \mathbf{H} - \frac{\partial \mathbf{P}}{\partial t}$$
(5a)

$$\frac{\partial \mathbf{H}}{\partial t} = -\frac{1}{\mu_0} \nabla \times \mathbf{E}$$
(5b)

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$$\ddot{\mathcal{P}} + 2\nu \dot{\mathcal{P}} + \omega_0^2 \mathcal{P} = \epsilon_0 \omega_\rho^2 \mathbf{E}$$
(5c)

with

$$\mathbf{P}(t,\mathbf{x}) = \int_{a}^{b} \mathcal{P}(t,\mathbf{x};\omega_{0}^{2})f(\omega_{0}^{2})d\omega_{0}^{2}.$$

## Theorem (Stability of Maxwell-Random Lorentz)

Let  $\mathcal{D} \subset \mathbb{R}^2$  and suppose that  $\mathbf{E} \in C(0, T; H_0(\operatorname{curl}, \mathcal{D})) \cap C^1(0, T; (L^2(\mathcal{D}))^2)$ ,  $\mathcal{P} \in C^1(0, T; (L^2(\Omega) \otimes L^2(\mathcal{D}))^2)$ , and  $H(t) \in C^1(0, T; L^2(\mathcal{D}))$  are solutions of the weak formulation for the Maxwell-Random Lorentz system along with PEC boundary conditions. Then the system exhibits energy decay

 $\mathcal{E}(t) \leq \mathcal{E}(0) \ \forall t \geq 0,$ 

where the energy  $\mathcal{E}(t)$  is defined as

$$\mathcal{E}(t) = \left( \left\| \sqrt{\mu_0} \ H(t) \right\|_2^2 + \left\| \sqrt{\epsilon_0 \epsilon_\infty} \ \mathbf{E}(t) \right\|_2^2 + \left\| \frac{\omega_0}{\omega_p \sqrt{\epsilon_0}} \ \mathcal{P}(t) \right\|_F^2 + \left\| \frac{1}{\omega_p \sqrt{\epsilon_0}} \ \mathcal{J}(t) \right\|_F^2 \right)^{\frac{1}{2}}$$
(6)  
where  $\| u \|_F^2 = \mathbb{E}[\| u \|_2^2]$  and  $\mathcal{J} := \frac{\partial \mathcal{P}}{\partial t}.$ 

Proof involves showing that

$$rac{d\mathcal{E}(t)}{dt} = rac{-1}{\mathcal{E}(t)} \Big\| \sqrt{rac{2
u}{\epsilon_0 \omega_
ho^2}} \mathcal{J} \Big\|_F^2 \leq 0.$$

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# **Polynomial Chaos**

We wish to approximate the random polarization with orthogonal polynomials of the standard random variable  $\xi$ . Let  $\omega_0^2 = r\xi + m$  and  $\xi \in [-1, 1]$ . Suppressing the dimension of P and the spatial dependence, we have

$$\mathcal{P}(\xi, t) = \sum_{i=0}^{\infty} \alpha_i(t) \phi_i(\xi) \rightarrow \ddot{\mathcal{P}} + 2\nu \dot{\mathcal{P}} + \omega_0^2 \mathcal{P} = \epsilon_0 \omega_p^2 E.$$

Utilizing the Triple Recursion Relation for orthogonal polynomials:

$$\xi \phi_n(\xi) = a_n \phi_{n+1}(\xi) + b_n \phi_n(\xi) + c_n \phi_{n-1}(\xi).$$

the differential equation becomes

$$\sum_{i=0}^{\infty} \left[\ddot{\alpha}_i(t) + 2\nu\dot{\alpha}_i(t) + m\alpha_i(t)\right]\phi_i(\xi) + r\sum_{i=0}^{\infty}\alpha_i(t)\left[\mathbf{a}_i\phi_{i+1}(\xi) + b_i\phi_i(\xi) + c_i\phi_{i-1}(\xi)\right] = \epsilon_0\omega_p^2 E\phi_0(\xi).$$

# Galerkin Projection

We apply a Galerkin Projection onto the space of polynomials of degree at most p:

$$\ddot{\vec{\alpha}} + 2\nu\dot{\vec{\alpha}} + A\vec{\alpha} = \vec{f}$$

$$A = rM + mI, \quad M = \begin{pmatrix} b_0 & c_1 & 0 & \cdots & 0\\ a_0 & b_1 & c_2 & & \vdots\\ 0 & \ddots & \ddots & \ddots & 0\\ \vdots & & a_{p-2} & b_{b-1} & c_p\\ 0 & \cdots & 0 & a_{p-1} & b_p \end{pmatrix}$$

Or we can write as a first order system:

$$\dot{\vec{\alpha}} = \vec{\beta} \\ \dot{\vec{\beta}} = -A\vec{\alpha} - 2\nu I\vec{\beta} + \vec{f},$$

where  $\vec{f} = \hat{e}_1 \epsilon_0 \overline{\omega}_p^2 E$  with  $\overline{\omega}_p$  meaning expected value.

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# Maxwell-PC Lorentz

The polynomial chaos system coupled with 1D Maxwell's equations becomes

$$\begin{aligned} \epsilon_{\infty}\epsilon_{0}\frac{\partial E}{\partial t} &= -\frac{\partial H}{\partial z} - \beta_{0} \\ \frac{\partial H}{\partial t} &= -\frac{1}{\mu_{0}}\frac{\partial E}{\partial z} \\ \dot{\vec{\alpha}} &= \vec{\beta} \\ \dot{\vec{\beta}} &= -A\vec{\alpha} - 2\nu I\vec{\beta} + \vec{f} \end{aligned}$$

Initial Conditions:

$$E(0,z) = H(0,z) = \vec{\alpha}(0,z) = \vec{\beta}(0,z) = 0$$

Boundary Conditions:

$$E(t,0) = E_L(t)$$
 and  $E(t,z_R) = 0$ 

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# **FDTD** Discretization





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We stagger three discrete meshes in the x and y directions and two discrete meshes in time:

$$\begin{split} \tau_h^{E_x} &:= \left\{ \left( x_{\ell + \frac{1}{2}}, y_j \right) | 0 \le \ell \le L - 1, 0 \le j \le J \right\} \\ \tau_h^{E_y} &:= \left\{ \left( x_{\ell}, y_{j + \frac{1}{2}} \right) | 0 \le \ell \le L, 0 \le j \le J - 1 \right\} \\ \tau_h^{H} &:= \left\{ \left( x_{\ell + \frac{1}{2}}, y_{j + \frac{1}{2}} \right) | 0 \le \ell \le L - 1, 0 \le j \le J - 1 \right\} \\ \tau_t^{E} &:= \left\{ (t^n) | 0 \le n \le N \right\} \\ \tau_t^{H} &:= \left\{ \left( t^{n + \frac{1}{2}} \right) | 0 \le n \le N - 1 \right\}. \end{split}$$

Let U be one of the field variables: H,  $E_x$ ,  $E_y$ ,  $\vec{\alpha}_x$ ,  $\vec{\alpha}_y$ ,  $\vec{\beta}_x$ . Let  $(x_i, y_j)$  be a node on any discrete spatial mesh, and  $\gamma$  be either n or  $n + \frac{1}{2}$  with  $\gamma \leq N$ . We define the *grid functions* or the numerical approximations

$$U_{i,j}^{\gamma} \approx U(x_i, y_j, t^{\gamma}).$$

We define the centered temporal difference operator and a discrete time averaging operation as

$$\delta_{t} U_{i,j}^{\gamma} := \frac{U_{i,j}^{\gamma+\frac{1}{2}} - U_{i,j}^{\gamma-\frac{1}{2}}}{\Delta t}, \qquad (9)$$
$$\overline{U}_{i,j}^{\gamma} := \frac{U_{i,j}^{\gamma+\frac{1}{2}} + U_{i,j}^{\gamma-\frac{1}{2}}}{2}, \qquad (10)$$

and the centered spatial difference operators in the x and y direction, respectively as

$$\delta_{x} U_{i,j}^{\gamma} := \frac{U_{i+\frac{1}{2},j}^{\gamma} - U_{i-\frac{1}{2},j}^{\gamma}}{\Delta x},$$
(11)  
$$\delta_{y} U_{i,j}^{\gamma} := \frac{U_{i,j+\frac{1}{2}}^{\gamma} - U_{i,j-\frac{1}{2}}^{\gamma}}{\Delta y}.$$
(12)

# Maxwell-PC Lorentz-FDTD

The Yee Scheme applied to the Maxwell-PC Lorentz yields

$$\delta_t H^n_{\ell+\frac{1}{2},j+\frac{1}{2}} = \frac{1}{\mu_0} \left( \delta_y E^n_{x_{\ell+\frac{1}{2},j+\frac{1}{2}}} - \delta_x E^n_{y_{\ell+\frac{1}{2},j+\frac{1}{2}}} \right)$$
(13a)

$$\epsilon_{0}\epsilon_{\infty}\delta_{t}E_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}} = \delta_{y}H_{\ell+\frac{1}{2},j}^{n+\frac{1}{2}} - \overline{\beta}_{0,x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}}$$
(13b)

$$\epsilon_{0}\epsilon_{\infty}\delta_{t}E_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}} = -\delta_{x}H_{\ell,j+\frac{1}{2}}^{n+\frac{1}{2}} - \overline{\beta}_{0,y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}}$$
(13c)

$$\delta_t \vec{\alpha}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}} = \overline{\vec{\beta}}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}}$$
(13d)

$$\delta_t \vec{\alpha}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}} = \overline{\vec{\beta}}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}}$$
(13e)

$$\delta_t \vec{\beta}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}} = -A \vec{\alpha}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}} - 2\nu \vec{\beta}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}} + \hat{e}_1 \epsilon_0 \omega_\rho^2 \vec{E}_{x_{\ell+\frac{1}{2},j}}^{n+\frac{1}{2}}$$
(13f)

$$\delta_t \vec{\beta}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}} = -A \vec{\alpha}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}} - 2\nu \vec{\beta}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}} + \hat{e}_1 \epsilon_0 \omega_p^2 \overline{E}_{y_{\ell,j+\frac{1}{2}}}^{n+\frac{1}{2}}.$$
(13g)

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#### Stability

# Staggered $L^2$ normed spaces

Next, we define the  $L^2$  normed spaces

$$\mathbb{V}_{\mathcal{E}} := \left\{ \mathbf{F} : \tau_{h}^{\mathcal{E}_{X}} \times \tau_{h}^{\mathcal{E}_{Y}} \longrightarrow \mathbb{R}^{2} \mid \mathbf{F} = (F_{x_{l+\frac{1}{2},j}}, F_{y_{l,j+\frac{1}{2}}})^{T}, \|\mathbf{F}\|_{\mathcal{E}} < \infty \right\}$$
(14)

$$\mathbb{V}_{H} := \left\{ U : \tau_{h}^{H} \longrightarrow \mathbb{R} \mid U = (U_{l+\frac{1}{2}, j+\frac{1}{2}}), \|U\|_{H} < \infty \right\}$$
(15)

with the following discrete norms and inner products

$$\|\mathbf{F}\|_{E}^{2} = \Delta \times \Delta y \sum_{\ell=0}^{L-1} \sum_{j=0}^{J-1} \left( |F_{x_{\ell+\frac{1}{2},j}}|^{2} + |F_{y_{\ell,j+\frac{1}{2}}}|^{2} \right), \forall \mathbf{F} \in \mathbb{V}_{E}$$
(16)

$$(\mathbf{F}, \mathbf{G})_{E} = \Delta x \Delta y \sum_{\ell=0}^{L-1} \sum_{j=0}^{J-1} \left( F_{x_{\ell+\frac{1}{2},j}} G_{x_{\ell+\frac{1}{2},j}} + F_{y_{\ell,j+\frac{1}{2}}} G_{y_{\ell,j+\frac{1}{2}}} \right), \forall \mathbf{F}, \mathbf{G} \in \mathbb{V}_{E}$$
(17)

$$\|U\|_{H}^{2} = \Delta \times \Delta y \sum_{\ell=0}^{L-1} \sum_{j=0}^{J-1} |U_{\ell+\frac{1}{2},j+\frac{1}{2}}|^{2}, \forall \ U \in \mathbb{V}_{H}$$
(18)

$$(U,V)_{H} = \Delta x \Delta y \sum_{\ell=0}^{L-1} \sum_{j=0}^{J-1} U_{\ell+\frac{1}{2},j+\frac{1}{2}} V_{\ell+\frac{1}{2},j+\frac{1}{2}}, \forall U, V \in \mathbb{V}_{H}.$$
 (19)

#### Stability

We define a space and inner product for the random polarization in vector notation, since  $\vec{\alpha}$  and  $\vec{\beta}$  are now  $2 \times p + 1$  matrices:

$$\mathbb{V}_{\alpha} := \left\{ \vec{\boldsymbol{\alpha}} : \tau_{h}^{\mathcal{E}_{x}} \times \tau_{h}^{\mathcal{E}_{y}} \longrightarrow \mathbb{R}^{2} \times \mathbb{R}^{p+1} \ \Big| \ \vec{\boldsymbol{\alpha}} = [\boldsymbol{\alpha}_{0}, \dots, \boldsymbol{\alpha}_{p}], \boldsymbol{\alpha}_{k} \in \mathbb{V}_{\mathcal{E}}, \|\vec{\boldsymbol{\alpha}}\|_{\alpha} < \infty \right\}$$

where the discrete  $L^2$  grid norm and inner product are defined as

$$\begin{split} \|ec{lpha}\|_{lpha}^2 &= \sum_{k=0}^p \|lpha_k\|_E^2, \quad orall ec{lpha} \in \mathbb{V}_{lpha} \ (ec{lpha},ec{eta})_{lpha} &= \sum_{k=0}^p \left( lpha_k, eta_k 
ight)_E, \quad orall ec{lpha}, ec{eta} \in \mathbb{V}_{lpha}. \end{split}$$

We choose both spatial steps to be uniform and equal  $(\Delta x = \Delta y = h)$ , and require that the usual CFL condition for two dimensions holds:

$$\sqrt{2}c_{\infty}\Delta t \le h.$$
 (20)

#### Stability

# Theorem (Energy Decay for Maxwell-PC Lorentz-FDTD)

If the stability condition (20) is satisfied, then the Yee scheme for the 2D TE mode Maxwell-PC Lorentz system given in (13) satisfies the discrete identity

$$\delta_t \mathcal{E}_h^{n+\frac{1}{2}} = \frac{-1}{\overline{\mathcal{E}}_h^{n+\frac{1}{2}}} \left\| \sqrt{\frac{2\nu}{\epsilon_0 \omega_\rho^2}} \overline{\vec{\beta}}_h^{n+\frac{1}{2}} \right\|_A^2 \tag{21}$$

for all n where

$$\mathcal{E}_{h}^{n} = \left(\mu_{0}(H^{n+\frac{1}{2}}, H^{n-\frac{1}{2}})_{H} + \left\|\sqrt{\epsilon_{0}\epsilon_{\infty}} \mathbf{E}^{n}\right\|_{E}^{2} + \left\|\sqrt{\frac{\omega_{0}^{2}}{\epsilon_{0}\omega_{p}^{2}}}\vec{\alpha}^{n}\right\|_{\alpha}^{2} + \left\|\sqrt{\frac{1}{\epsilon_{0}\omega_{p}^{2}}}\vec{\beta}^{n}\right\|_{\alpha}^{2}\right)^{1/2}$$

$$defines a discrete energy.$$
(22)

In the above  $\|\vec{\alpha}\|_A^2 := (A\vec{\alpha}, \vec{\alpha})_{\alpha}$  given A positive definite, which is true iff r < m. Note that  $\|\vec{\alpha}\|_{\alpha}^2 \approx \|\mathbb{E}[\mathcal{P}]\|_2^2 + \|\text{StdDev}(\mathcal{P})\|_2^2 = \mathbb{E}[\|\mathcal{P}\|_2^2] = \|\mathcal{P}\|_F^2$  so that this is a natural extension of the Maxwell-Random Lorentz energy (6).

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#### Theorem

The dispersion relation for the Maxwell-Random Lorentz system is given by

$$\frac{\omega^2}{c^2} \epsilon(\omega) = \|\mathbf{k}\|^2$$

where the expected complex permittivity is given by

$$\epsilon(\omega) = \epsilon_{\infty} + (\epsilon_s - \epsilon_{\infty}) \mathbb{E}\left[\frac{\omega_0^2}{\omega_0^2 - \omega^2 - i2\nu\omega}\right].$$

Where **k** is the wave vector and  $c = 1/\sqrt{\mu_0 \epsilon_0}$  is the speed of light.

The exact dispersion relation can be compared with a discrete dispersion relation to determine the amount of dispersion error.

We define the phase error  $\Phi$  for a scheme applied to a model to be

$$\Phi = \left| \frac{k_{EX} - k_{\Delta}}{k_{EX}} \right|,\tag{23}$$

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where the numerical wave number  $k_{\Delta}$  is implicitly determined by the corresponding discrete dispersion relation and  $k_{EX}$  is the exact wave number for the given model.

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where the numerical wave number  $k_{\Delta}$  is implicitly determined by the corresponding discrete dispersion relation and  $k_{EX}$  is the exact wave number for the given model.

- We wish to examine the phase error as a function of ω in the range around ω
  <sub>0</sub>. Δt is determined by h := ω
  <sub>0</sub>Δt/(2π), while Δx = Δy are determined by the CFL condition.
- We assume a uniform distribution and the following parameters Lorentz material:

$$\epsilon_{\infty}=1, \quad \epsilon_s=2.25, \quad \nu=2.8\times 10^{15} \ 1/\text{sec}, \quad \overline{\omega}_0=4\times 10^{16} \ \text{rad/sec}.$$



Figure 4: Plots of phase error at  $\theta = 0$ .

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Figure 5: Plots of phase error at  $\theta = 0$ .



Figure 6: Plots of phase error at  $\theta = 0$ .

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Figure 7: Plots of phase error at  $\theta = 0$ .



Figure 8: Plots of phase error at  $\theta = 0$ .



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Current Work:

• Analyze the dispersion error of the Random Lorentz model Future Work:

- Extend to nonlinear polarization models
- Allow  $\epsilon_s$ ,  $\epsilon_\infty$  to be uncertain.

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